

From observation to publication

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Abstract

This article describes the use of \TeX in publishing observations of variable stars observed by Dutch amateur-astronomers. The observations are published in the journal "Variabilia" and in the so-called Reports. In the latter the observations, collected in several years, are published and submitted to the professional astronomer. It includes tables and light-curves: plot of the changing magnitude of the star versus time. In creating the light-curves: \PCTeX is used. In preparing the files for \PCTeX simple \TeX -coding is used for manipulating the data.

1 The data

Each observation is characterized by the observed star, a date, the brightness of the star and a code representing the observer. To avoid calendar problems the so-called Julian Date is used. For example the Julian Date of 1961, september 9, my date of birth, is equal to 2437552.

Each observation is \TeX - coded as followed:

```
\obs #1.#2.#3.#4.#5.
```

#1 is the integer of the Julian Date, #2 it's fraction. #3 is the integer of the magnitude, #4 it's fraction and #5 is the observer-code. The latter is not used for plotting. \TeX can calculate with integers so the observing date and magnitude, the apparent brightness of a star, are transformed to integers. In case of date it is simply the integer of the Julian date. The magnitude is expressed in units of 0.1 so the magnitude is $10 \times \#3 + \#4$.

For each star we have a number of observations. The file of observations processed by plain- \TeX produces a file suitable for \PCTeX . This .plt file contains all the necessary information for a plot: the axis are proper scaled and labeled and contains information about the star. See the example at page 119 Figure 1 shows an typical file ready for \TeX .

The macros used to create the .plt file are given below. Part II neglect the fraction of the Julian Date and calculates the magnitude, in units of 0.1. Both are written to a temporary file with the extension .obs. In part II also the minimum and maximum values along the axis are estimated. These values are used in Part III to calculate the proper sizes of the graph. In Part III the final .plt file is created and can be processed by \PCTeX .

```
%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%
%% PART I plot.tex %%
%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%
\newcount\tempx \newcount\tempy
\newcount\tempz \newcount\maxx
\newcount\minx \newcount\miny
\newcount\maxy \newcount\numobs
\newwrite\plotfile \newwrite\obsfile
\def\head#1{
\global\def\plotname{\sternum\ #1}
\initplot}
\def\endhead{\immediate\closeout\obsfile
\startplot}
\def\harvard#1{\global\def\sternum{#1}}
\def\type#1{\global\def\typename{#1}}
\def\initplot{
\immediate\openout\obsfile=\jobname.obs
\global\minx=99999
\global\maxx=0
\global\miny=99999
\global\maxy=0
\global\numobs=0}
\def\pplot#1#2{\tempx=#1
\tempy=\maxy
\tempz=#2
\advance\tempy by-\tempz
\advance\tempx by-\minx
\immediate\write\plotfile{\noexpand
\put {$\noexpand\bullet$} at
{\the\tempx} {\the\tempy}}}
```

```

%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%
%% PART II %%
%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%
\def\obs #1.#2.#3.#4.#5.{
\message{#5}
\global\advance\numobs by 1
\tempx=#1
% estimate min max x
\ifnum\minx>\tempx \global\minx=\tempx \fi
\ifnum\maxx<\tempx \global\maxx=\tempx \fi
\tempy=#3
\tempz=#4
% calculating magnitude
\multiply\tempy by 10
\advance\tempy by\tempz
% estimate min max y
\ifnum\miny>\tempy \global\miny=\tempy \fi
\ifnum\maxy<\tempy \global\maxy=\tempy \fi
\immediate\write\obsfile{\noexpand
\pplot{\the\tempx}{\the\tempy}}

%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%
%% PART III %%
%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%
\def\startplot{
% create nice numbers along axis
\global\divide\minx by 10
\global\multiply\minx by 10
\global\divide\maxx by 10
\global\multiply\maxx by 10
\global\advance\maxx by 10
\global\divide\miny by 10
\global\multiply\miny by 10
\global\divide\maxy by 10
\global\multiply\maxy by 10
\global\advance\maxy by 10
\global\advance\maxx by-\minx
\ifnum\numobs>15
% writing pictex commands
\def\name{\sternum.plt}
\immediate\openout\plotfile=\name
\message{\name}
\immediate\write\plotfile{\noexpand
\beginpicture}
\immediate\write\plotfile{\noexpand
\setcoordinatesystem units <0.35mm,1mm> }
\tempz=\maxy
\advance\tempz by-\miny
\immediate\write\plotfile{\noexpand
\setplotarea x from 0 to {\the\maxx},
y from 0 to {\the\tempz}}
\immediate\write\plotfile{\noexpand
\plothead{\plotname\ \typename}}
\immediate\write\plotfile{\noexpand
\axis bottom label {JD-{\the\minx}}}
\immediate\write\plotfile{ticks
numbered from 0 to {\the\maxx} by 40 / }
\immediate\write\plotfile{\noexpand
\axis left label { } }

\immediate\write\plotfile{ticks
from 0 to {\the\tempz} by 10 / }
\tempx=\miny
\advance\tempx by-10
\tempz=\maxy
\loop\ifnum\tempz>\tempx
\tempy=\maxy
\advance\tempy by-\tempz
\immediate\write\plotfile{\noexpand
\put {\the\tempz} at
-15 {\the\tempy}}
\advance\tempz by-10
\repeat
% find out where to put text
% along y-axis.
\tempz=\maxy
\advance\tempz by-\miny
\divide\tempz by2
\tempx=\tempz
\divide\tempx by10
\tempy=\tempz
\multiply\tempx by10
\advance\tempy by-\tempx
\ifnum\tempy=0
\advance\tempz by5
\fi
\immediate\write\plotfile{\noexpand
\put {Magn.} at -15 {\the\tempz}}
\input \jobname.obs
\immediate\write\plotfile{\noexpand\endpicture}
\immediate\closeout\plotfile
\fi}

\input plot.tex
\harvard{000451}
\head{SS Cas}
\type{Mira HIP}
\obs 47537.4.10.0.CMG.
\obs 47539.4.9.7.FJH.
\obs 47544.3.10.2.CMG.
\obs 47544.3.10.4.BMU.
\obs 47549.4.10.5.BMU.
\obs 47550.4.10.3.CMG.
\obs 47554.4.10.5.NWL.
\obs 47565.3.11.3.CMG.
\obs 47565.3.11.4.BMU.
\obs 47567.3.11.5.FJH.
\obs 47569.3.11.7.BMU.
\obs 47573.3.11.7.CMG.
\obs 47574.4.12.0.BMU.
\obs 47578.3.12.1.CMG.
\obs 47579.3.11.8.FJH.
\obs 47579.3.12.1.BMU.
\obs 47586.4.12.3.BMU.
\obs 47592.3.12.6.CMG.
\obs 47594.3.12.5.FJH.
\endhead
\bye

```

Fig. 1: An input-file for \TeX to create a file ready for $\text{P}\text{f}\text{C}\text{I}\text{E}\text{X}$.

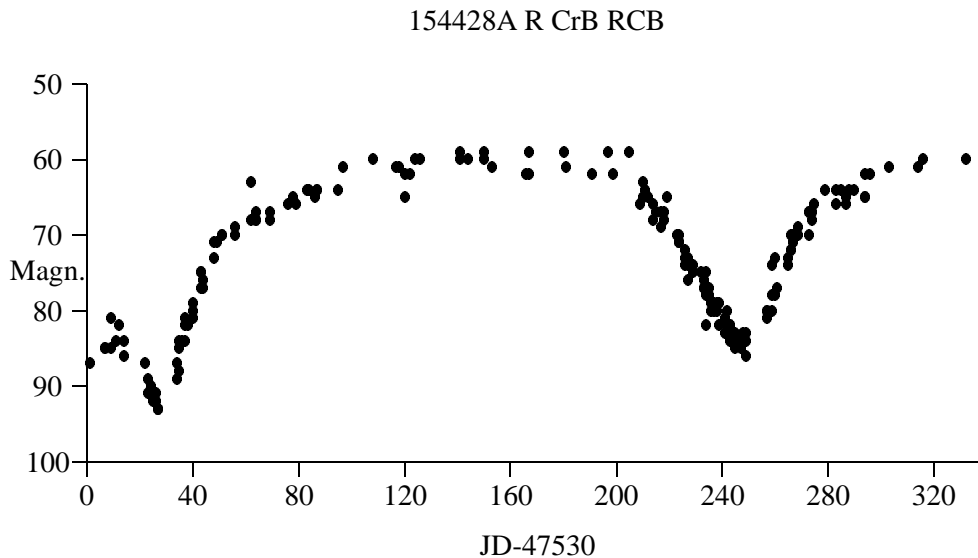


Fig. 2: The light-curve of the star R CrB. The decrease in brightness can be seen clearly in this picture. The data are collected by Dutch amateur-astronomers.

2 Macro's

In Part I the initialization is done: counters and files are defined. Information about the stars are stored in the tokens `\plotname`, `\sternum` and `\typename`. Also some initial values for the minimum and maximum values along the axis are set. The macro `\pplot` is used in Part II and executed in Part III. The macro writes the re-scaled X and Y values to the `.plt` file. Again the latter is used with \LaTeX to obtain the desired result.

In Part II `\obs` is defined: the magnitude and Julian Date are converted to integers as described above and the maximum and minimum values are estimated according to the simple algorithm:

```
if  $x_i < \min$  then  $\min := x_i$ ;
if  $x_i > \max$  then  $\max := x_i$ ;
```

Also the number of observations is counted. This number is stored in `\numobs`.

In Part III the final plotting commands are written to the `.plt` file. This part is only executed if the number of observations is bigger than 15. The minima and maxima found in Part II are re-calculated to have nice numbers along the axis. Commands for labels and numbering axis are also written into the file. After setting up the graph the data to be plotted are read from the `.obs` file. This file is created in Part II. Figure 3 shows such a `.plt` file.

```
\beginpicture
\setcoordinatesystem units <0.35mm,1mm>
```

```
\setplotarea x from 0 to {340},
y from 0 to {50}
\plotheadings {154428A\ R CrB\ RCB}
\axis bottom label {JD-{47530}}
ticks numbered from 0 to {340} by 40 /
\axis left label {}
ticks from 0 to {50} by 10 /
\put {100} at -15 {0}
\put {90} at -15 {10}
\put {80} at -15 {20}
\put {70} at -15 {30}
\put {60} at -15 {40}
\put {50} at -15 {50}
\put {Magn.} at -15 {25}
\put {$\bullet$} at {1} {13}
\put {$\bullet$} at {7} {15}
:
\endpicture
```

Fig. 3: The start of a typical plot-file as created by Part III.

3 Publishing of the data

All the observations of one year are collected in one file with a structure as in Figure 1. This file contains 6000 to 10.000 lines depending on the weather conditions. The file is processed using `\plot.tex` and a number of `.plt` files are created. The same file is used to create a six-column tabular output of the observations, see the example at the next page. These macro's are available on request.

okt – nov – dec 1991

000451 SS Cas Mira HIP	004533 RR And Mira	011208 S Psc Mira	013937 AR And UGSS	021281 Z Cep Mira
560.4 13.0 FJH	533.5 15.0 FJH	540.5 11.0 FJH	594.45 11.9 FJH	536.5 14.1 FJH
572.5 13.3 FJH	557.3 13.8 FJH	557.5 11.8 FJH	596.32 12.2 FJH	551.4 13.5 FJH
594.5 12.5 FJH	569.5 12.1 FJH	570.5 12.2 FJH	601.46 12.7 FJH	570.5 11.9 FJH
	596.4 10.9 FJH	596.3 12.7 FJH		581.3 11.6 FJH
000928 UW And Mira	611.4 9.8 FJH		015254 U Per Mira HIP	590.6 11.8 FJH
533.4 11.7 FJH	004746 A RV Cas Mira	011712 U Psc Mira	533.4 8.3 KKP	613.5 12.3 FJH
556.3 12.7 FJH	533.5 15.5 FJH	536.4 14.2 FJH	536.4 8.5 JOJ	0214-0 3 Mira Mira HIP
570.5 13.7 FJH	557.3 15.0 FJH	558.4 14.4 FJH	556.3 8.2 JOJ	
	601.5 12.6 FJH	570.5 13.6 FJH	556.6 8.7 KKP	536.5 4.0 SAQ
001046 X And Mira	611.4 11.7 FJH	596.3 12.2 FJH	601.3 9.1 KKP	536.6 4.0 BMU
532.4 10.9 FJH			601.4 8.4 JOJ	596.3 6.1 SAQ
551.4 9.9 FJH	004958 W Cas Mira HIP	012020 RX Psc Mira	015457 V666 Cas Mira	021558 S Per SRc HIP
572.4 9.1 FJH	556.6 :10.1 KKP	536.4 :15.1 FJH	556.4 11.2 FJH	556.4 12.0 FJH
596.4 9.1 FJH		596.3 14.5 FJH	576.3 11.2 FJH	557.3 11.8 HIL
611.4 9.8 FJH	005840 RX And UGZ	012031 TY Psc UGSU		576.3 12.2 FJH
	532.42 14.0 FJH	594.45 12.1 FJH	015912 S Ari Mira	594.5 12.2 FJH
001726 T And Mira	533.48 13.7 FJH	596.34 12.4 FJH	533.5 13.8 FJH	0220-0 0 R Cet Mira HIP
557.4 10.6 FJH	536.43 14.0 FJH		558.4 14.5 FJH	
572.4 9.5 FJH	540.44 11.5 FJH	012502 R Psc Mira	570.5 14.6 FJH	536.5 9.2 SAQ
596.4 8.2 FJH	545.38 12.0 FJH	536.5 13.8 FJH	596.3 :15.2 FJH	601.3 8.3 SAQ
614.4 8.5 FJH	551.34 13.5 FJH	559.3 13.9 FJH		
	556.35 13.0 FJH	570.5 13.9 FJH	020227 Z Tri Mira	022150 RR Per Mira
001755 T Cas Mira HIP	557.32 13.4 FJH	596.3 13.1 FJH	532.5 14.7 FJH	540.5 11.5 FJH
596.4 9.0 FJH	558.37 13.1 FJH		540.5 14.7 FJH	556.4 11.7 FJH
614.4 9.1 FJH	559.38 13.5 FJH	012746 SX And Mira	596.3 13.7 FJH	575.4 12.2 FJH
	560.35 13.6 FJH	556.4 10.6 FJH		590.5 12.7 FJH
001838 R And Mira HIP	569.46 12.3 FJH	575.3 11.1 FJH	020356 UV Per UGSS	022980 RR Cep Mira
540.5 9.1 FJH	570.35 11.2 FJH	594.5 11.5 FJH	613.51 12.6 FJH	536.5 14.1 FJH
557.3 9.8 FJH	572.39 11.5 FJH	611.4 12.0 FJH	614.40 12.7 FJH	570.5 13.1 FJH
572.4 10.3 FJH	574.43 11.6 FJH		615.23 12.7 FJH	581.3 12.5 FJH
596.3 10.9 FJH	575.34 11.6 FJH	013050 KT Per UGZ		590.6 11.8 FJH
611.4 11.2 FJH	576.33 12.2 FJH	536.44 12.4 FJH	020657 A TZ Per UGZ	613.5 11.2 FJH
	594.44 13.7 FJH	540.46 14.4 FJH	532.50 12.9 FJH	
002725 A TU And Mira HIP	596.35 13.7 FJH	556.35 12.2 FJH	536.44 13.8 FJH	023133 R Tri Mira HIP
557.4 11.5 FJH	615.28 13.9 FJH	557.32 11.9 FJH	540.46 14.3 FJH	533.4 9.7 KKP
	010621 A X Psc Mira	558.36 11.8 FJH	551.35 12.8 FJH	540.3 10.6 HIL
003162 TY Cas Mira	536.4 14.5 FJH	559.38 11.8 FJH	557.33 13.9 FJH	557.3 11.6 HIL
576.3 13.5 FJH	558.4 13.9 FJH	560.35 12.0 FJH	558.36 14.2 FJH	557.5 11.8 FJH
594.5 11.5 FJH	570.5 13.5 FJH	572.46 14.4 FJH	559.39 14.2 FJH	570.5 12.0 FJH
	596.4 12.9 FJH	575.35 12.4 FJH	560.48 13.7 FJH	594.5 12.3 FJH
003179 Y Cep Mira	010937 FO And UG	576.33 12.4 FJH	570.35 13.5 FJH	
551.4 13.0 FJH	533.47 14.9 FJH	590.52 11.9 FJH	572.47 13.8 FJH	030226 Z Ari Mira
570.5 13.6 FJH	536.45 <15.4 FJH	601.46 11.9 FJH	575.36 13.7 FJH	533.5 13.8 FJH
590.6 14.0 FJH	572.46 14.6 FJH		590.51 13.6 FJH	559.4 :14.4 FJH
	596.34 <15.4 FJH	013238 RU And SRa	596.46 13.6 FJH	
004047 U Cas Mira	615.28 14.2 FJH	536.5 12.0 FJH	601.46 12.7 FJH	030514 U Ari Mira
540.4 12.7 FJH		556.4 11.7 FJH	615.27 13.4 FJH	
550.4 13.0 FJH	010940 U And Mira	575.3 11.9 FJH		540.5 10.5 FJH
557.3 13.7 FJH	540.5 9.6 FJH		021024 R Ari Mira HIP	559.4 11.6 FJH
572.5 14.2 FJH	557.4 10.2 FJH	013338 Y And Mira	551.4 12.4 FJH	596.3 13.6 FJH
596.3 :15.5 FJH	569.5 11.1 FJH	532.4 14.2 FJH	575.5 10.5 FJH	
	594.4 11.9 FJH	540.5 14.0 FJH		031170 V667 Cas Mira
004132 RW And Mira	614.4 12.3 FJH	551.3 13.2 FJH	021143 A W And Mira HIP	572.5 10.5 HIL
533.5 15.3 FJH		569.5 11.3 FJH	540.4 13.1 FJH	596.5 9.4 FJH
596.3 :15.5 FJH	011041 A UZ And Mira	594.5 9.9 FJH	551.3 12.7 FJH	600.5 9.8 HIL
	533.5 15.1 FJH	611.4 9.2 FJH	569.5 12.4 FJH	
004435 V And Mira	557.3 14.0 FJH		594.5 11.8 FJH	032043 Y Per Mira
557.3 14.1 FJH	569.5 13.9 FJH	013937 AR And UGSS	611.4 10.7 FJH	533.4 9.1 KKP
569.5 14.5 FJH	594.4 13.2 FJH	532.42 14.2 FJH		556.6 9.8 KKP
596.3 14.8 FJH	614.4 12.0 FJH	533.47 15.3 FJH		580.5 10.1 KKP
615.3 14.3 FJH		556.36 12.5 FJH		584.5 10.3 KKP
	011055 A VZ Cas Mira	557.32 12.8 FJH		597.4 9.7 KKP
	560.4 10.6 FJH	558.36 13.5 FJH		
	611.4 11.0 FJH	559.39 <15.0 FJH		
		572.36 12.3 FJH		
		574.43 12.7 FJH		
		575.36 14.0 FJH		

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